

## ARO EQUIPMENT SUMMARY AND STATUS

**12M Telescope** (EIE ALMA Prototype Antenna) – Kitt Peak July 30 2019



12 Meter Diameter; f/8 Observing season: October – June Carbon fiber construction Altitude over Azimuth Absolute pointing accuracy: 2.5" rms Tracking precision: 0.2" rms Efficiency: 85% +/- 5% Azimuth range: +/-270° Elevation range: 3° to 110° Observational range: ~18°- 84° Slew speed: 360° min<sup>-1</sup> Observational value: 60° min<sup>-1</sup> (Dome tracking limited) Supports position-switching, beamswitching, on-the-fly mapping, and continuum observing

#### **Front Ends**

Multi-Band Receiver (4mm, 3mm, 2mm, 1mm, Dual polarization, SBS)
3mm (ALMA Band 3; 84 – 116 GHz; Available)
Image Rejection: 15 dB minimum≥ 20 dB typical

Typical  $T_{sys@90GHz}$ = 120–160 K,  $T_{sys@115GHz}$ = 250–375 K Best  $T_{sys@90GHz}$ = 85–110 K,  $T_{sys@115GHz}$ = 150–200 K

4mm (66 – 90 GHz; Available, shared-risk)

1mm (211 – 275 GHz; Available, shared-risk)

2mm (In development)

#### **Back Ends**

#### ARO Wideband Spectrometer (AROWS)

Interim: two 4 GHz wide IFs (Available) Final: four 4 GHz wide IFs (Unavailable)

Mode	Total Bandwidth (MHz)	Channel Spacing (kHz)	Velocity Resolution at 86 GHz (km s <sup>-1</sup> )	Velocity Resolution at 115 GHz (km s <sup>-1</sup> )
0	4000	625	4.4	3.2
1	2000	312.5	2.2	1.6
2	1000	156.25	1.1	0.8
3	500	78.125	0.54	0.40
4	250	39.0625	0.28	0.20
5	125	19.53	0.14	0.10

#### **Observing Restrictions**

Sun Avoidance: The 12m requires a  $10^\circ$  Sun avoidance zone.

Observing: On-the-fly mapping is currently unavailable with the MAC back end.

#### **12 Meter MAC** (2 IF Mode; Available) Center frequency fixed at 6.3 GHz in

Center frequency fixed at 6.3 GHz in receiver IF.

Bandwidth (MHz)	Channels	Usable (MHz)	Usable Channels	$\triangle v^*$ (kHz)	Resolution (kHz)
800	2048	600	1536	390.6	781.2
800	4096	600	3072	195.3	390.6
400	4096	300	3072	97.6	195.3
400	8192	300	6144	48.8	97.6
200	8192	150	6144	24.4	48.8
200	16384	150	12288	12.2	24.4
100	16384	75	12288	6.1	12.2
100	32768	75	24576	3	6.1

\*This is the frequency sampling interval, not the FWHM channel width.

Astronomy

For more information, please contact:



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& Steward Observatorv



# ARO EQUIPMENT SUMMARY AND STATUS

f/13.8

Submillimeter Telescope (SMT) – Mt. Graham

construction Altitude over Azimuth

10 Meter Diameter, bent Cassegrain;

Observing season: October – June Carbon fiber and invar steel

Absolute pointing accuracy: 2" rms

Observational range: ~20° – 84°

Supports position-switching, beamswitching, on-the-fly mapping, and

Tracking precision: 0.2" rms Efficiency: 71% +/- 5% at 216 GHz and 228 GHz Azimuth range: +/-270° Elevation range: -2° to 91°

Slew speed: 60° min<sup>-1</sup>

continuum observing

July 30 2019



### **Front Ends**

1.3mm Receiver (Available) ALMA Band 6 (205 – 280 GHz) Dual polarization, SBS Typical T<sub>sys@230GHz</sub>= 200-275 K Best T<sub>sys@230GHz</sub>= 130-160 K

0.8mm Receiver (Available) 325 – 370 GHz Dual polarization, DSB Typical T<sub>sys@345GHz</sub>= 600–900 K Best T<sub>sys@345GHz</sub>= 350–500 K 0.4mm Receiver (Unavailable) 602 – 720 GHz Dual polarization, DSB 80 K T<sub>Rx</sub> 0.7mm Receiver (Unavailable) 385 – 500 GHz Dual polarization, SBS 150 K T<sub>Rx</sub>

New 0.8mm Band 7 Receiver (In development) ALMA Band 7 (275 – 373 GHz) Dual polarization, SBS

### **Back Ends**

### SMT Filter banks (Available)

The SMT IF range is 4–8 GHz for all receivers except the 0.8mm receiver, which has an IF range of 4–6 GHz. The IF center frequency is tunable from 4.5–7.5 GHz.

Mode	Bandwidth (MHz)	Resolution (kHz)
1 IF	2000	1000
2 IF	1000	1000
2 IF	256	250
4 IF	512	1000
4 IF	128	250

### **Observing Restrictions**

Sun Avoidance: The SMT requires a  $45^{\circ}\,\text{Sun}$  avoidance zone.

Observing: None.

For more information, please contact:





## ARO EQUIPMENT SUMMARY AND STATUS

Recently Fixed and Current Issues July 30 2019

**12M Telescope** (EIE ALMA Prototype Antenna) – Kitt Peak

MAC: Unusable for on-the-fly mapping due to hardware limitations.

### Submillimeter Telescope (SMT) – Mt. Graham

**Filters 1.0MHz**: There are periods of instability in the IF down conversion stages that manifests itself in platforming. Problems appear worse in position–switching and on–the–fly observing modes.

**Filters 250kHz**: The instability noted in the 1.0 MHz filters also appears in these filters. This platforming step is located at the IF center frequency and can be avoided by offsetting the IF.

**Position–switching mode:** There is a standing wave present when using observing in position–switching mode, especially at the upper end of 1.3mm receiver observing band. This standing wave is not atmospherically induced but instead is related to the receiver/hardware.

For more information, please contact: Dr. H. Alyson Ford, Manager, Radio Telescopes

